Lecture 6: The superstring

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STRING THEORY

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Motivation

There are several reasons to tackle superstring theory:

- Spectrum of the bosonic string has a tachyon; unstable (false?) vacuum.
- Nature has fermions.

We have to modify the (field content of the) world-sheet theory.

If we want to reproduce the Dirac equation for spacetime fermions,

$$i \Gamma^{\mu} p_{\mu} + m = 0$$

we must include anticommuting world-sheet fields, ψ^{μ} , with the expectation that their center of mass modes are Γ^{μ} , such that $\{\Gamma^{\mu}, \Gamma^{\nu}\} = 2\eta^{\mu\nu}$.

Notice that if this is the case, a kinetic term of the form

 $\psi^{\mu}\bar{\partial}\psi_{\mu} \quad \Rightarrow \quad \text{Cliff}_{D} \text{ algebra}$

Thus, D-dim Poincaré invariance makes the 2d world-sheet theory SUSY.

The superstring action

The equation of motion for spacetime spinors arises from the conserved supercurrent of the 2d theory.

Start with the gauge-fixed Polyakov action; add world-sheet Majorana spinors,

$$\boldsymbol{S} = -\frac{T}{2} \int_{\tau_1}^{\tau_2} d\tau \int_0^{\bar{\sigma}} d\sigma \left(\eta^{\alpha\beta} \partial_\alpha \boldsymbol{x}^\mu \, \partial_\beta \boldsymbol{x}_\mu - i \, \overline{\psi}^\mu \rho^\alpha \partial_\alpha \psi_\mu \right)$$

Here ρ^{α} are 2d Dirac matrices, $\{\rho^{\alpha}, \rho^{\beta}\} = -2\eta^{\alpha\beta}$; a convenient basis being,

$$\rho^{0} = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \qquad \rho^{1} = \begin{pmatrix} 0 & i \\ i & 0 \end{pmatrix}$$

We refer to the *real* ($\overline{\chi} = \chi^t \rho^0$, $\overline{\chi} \psi = \overline{\psi} \chi$) components of ψ in this basis as:

$$\psi = \left(\begin{array}{c} \psi_-\\ \psi_+ \end{array}\right) \qquad \psi_{\pm}^{\star} = \psi_{\pm}$$

It is important to emphasize that at this stage it is not obvious at all that such theory will lead to spacetime fermions or spacetime supersymmetry; in fact, the 2d fermion fields are bosons with respect to the spacetime Lorentz group.

Holomorphicity of Majorana-Weyl spinors

The fermionic kinetic term becomes

$$\overline{\psi} \cdot \rho^{\alpha} \partial_{\alpha} \psi = \psi_{-} \cdot \partial_{+} \psi_{-} + \psi_{+} \cdot \partial_{-} \psi_{+}$$

where we recall that the light-cone derivatives ∂_{\pm} are

$$\partial_{\pm} = \frac{\partial}{\partial \xi^{\pm}} = \frac{1}{2} \left(\partial_{\tau} \pm \partial_{\sigma} \right) \qquad \xi^{\pm} = \tau \pm \sigma$$

The equation of motion for the fermions is massless 2d Dirac equation,

$$\partial_+\psi_-^\mu=\partial_-\psi_+^\mu=\mathbf{0}$$

(still needs boundary conditions). Thus, the Majorana-Weyl fermions:

- $\psi^{\mu}_{-} = \psi^{\mu}_{-}(\xi^{-})$ describe right-movers while
- $\psi^{\mu}_{+} = \psi^{\mu}_{+}(\xi^{+})$ describe left-movers.

The equations of motion and constraints for x^{μ} are the same as before.

World-sheet supersymmetry

The system of free scalars and fermions is automatically invariant under the global, infinitesimal, world-sheet supersymmetry transformations

$$\delta_{\epsilon} \mathbf{X}^{\mu} = \overline{\epsilon} \, \psi^{\mu} \qquad \delta_{\epsilon} \psi^{\mu} = -i \, \rho^{\alpha} \partial_{\alpha} \mathbf{X}^{\mu} \, \epsilon$$

with ϵ a constant, anticommuting two-component spinor. Indeed,

$$\begin{split} \delta S &= -T \int_{\tau_1}^{\tau_2} d\tau \int_0^{\bar{\sigma}} d\sigma \left(\eta^{\alpha\beta} \partial_\alpha \delta_\epsilon x \cdot \partial_\beta x - \frac{i}{2} \left[\delta_\epsilon \overline{\psi} \cdot \rho^\alpha \partial_\alpha \psi + \overline{\psi} \cdot \rho^\alpha \partial_\alpha \delta_\epsilon \psi \right] \right) \\ &= -T \int_{\tau_1}^{\tau_2} d\tau \int_0^{\bar{\sigma}} d\sigma \left(\eta^{\alpha\beta} \bar{\epsilon} \partial_\alpha \psi + \frac{1}{2} \left[\bar{\epsilon} \rho^\beta \rho^\alpha \partial_\alpha \psi - \overline{\psi} \rho^\alpha \rho^\beta \epsilon \partial_\alpha \right] \right) \cdot \partial_\beta x \\ &= -T \int_{\tau_1}^{\tau_2} d\tau \int_0^{\bar{\sigma}} d\sigma \left[\left(\eta^{\alpha\beta} \bar{\epsilon} \partial_\alpha \psi + \frac{1}{2} \bar{\epsilon} \left\{ \rho^\beta, \rho^\alpha \right\} \partial_\alpha \psi \right) \cdot \partial_\beta x - \bar{\epsilon} \partial_\alpha \mathcal{J}^\alpha \right] \end{split}$$

where $\partial_{\alpha}\overline{\psi}\,\rho^{\alpha}\rho^{\beta}\epsilon = \overline{\epsilon}\,\rho^{\alpha}\rho^{\beta}\partial_{\alpha}\psi; \,\,\delta S$ is given by the Noether supercurrent

$$\delta S = T \int_{\tau_1}^{\tau_2} d\tau \int_0^{\bar{\sigma}} d\sigma \,\bar{\epsilon} \,\partial_\alpha \mathcal{J}^\alpha = 0 \qquad \mathcal{J}^\alpha = \frac{1}{2} \rho^\beta \rho^\alpha \psi \cdot \partial_\beta x$$

World-sheet supersymmetry

Some components of the supercurrent vanish. Indeed, due to $\rho^{\alpha}\rho^{\beta}\rho_{\alpha} = 0$,

 $\rho^{\alpha} \mathcal{J}_{\alpha} = \mathbf{0}$ only one chirality survives : \mathbf{J}_{α}

which in light-cone coordinates read

$$J_{\pm} = \psi_{\pm} \cdot \partial_{\pm} x$$

and they are conserved

 $\partial_{-}J_{+}=\partial_{+}J_{-}=0$

We can use the Noether method to compute the energy-momentum tensor,

$$T_{\pm\pm} = \partial_{\pm} x \cdot \partial_{\pm} x + \frac{i}{2} \psi_{\pm} \cdot \partial_{\pm} \psi_{\pm} \qquad T_{+-} = T_{-+} = 0$$

(the trace still vanishes). The Virasoro constraints $T_{++} = T_{--} = 0$ where used in the bosonic case to eliminate ghosts (negative norm states) coming from x^0 . What about those coming from ψ^0 ?

Boundary conditions and mode expansion: open strings

To remove the time-like component ψ^0 , we must also impose $J_+ = J_- = 0$. These come from the locally supersymmetric ungauged Polyakov action.

The mode decomposition and boundary conditions for $x^{\mu}(\tau, \sigma)$ are as before.

Let us consider the free fermionic fields $\psi^{\mu}(\tau, \sigma)$. In the case of open strings, the variational principle for the Polyakov action requires

$$\psi_+ \cdot \delta \psi_+ - \psi_- \cdot \delta \psi_- = \mathbf{0}$$
 at $\sigma = \mathbf{0}, \pi$

The Dirac equation admits two possible boundary conditions consistent with Lorentz invariance, namely

$$\psi^{\mu}_{\pm} = \pm \psi^{\mu}_{-}$$
 (and hence $\delta \psi^{\mu}_{\pm} = \pm \delta \psi^{\mu}_{-}$) at each end

The overall relative sign between ψ^{μ}_{-} and ψ^{μ}_{+} is irrelevant; we take, without loss of generality,

$$\psi^\mu_+(au, \mathbf{0}) = \psi^\mu_-(au, \mathbf{0})$$

This still leaves two possibilities at the other endpoint of the string, $\sigma = \pi$.

Boundary conditions and mode expansion: open strings

According to the relative sign,

 $\psi^{\mu}_{+}(\tau,\pi) = +\psi^{\mu}_{-}(\tau,\pi)$ Ramond (R) boundary condition $\psi^{\mu}_{+}(\tau,\pi) = -\psi^{\mu}_{-}(\tau,\pi)$ Neveu – Schwarz (NS) boundary condition

The mode expansion of the Dirac equation becomes

$$\psi_{\pm}^{\mu}(\tau \pm \sigma) = \frac{1}{\sqrt{2}} \sum_{r \in \mathbb{Z}} \psi_{r}^{\mu} e^{-ir(\tau \pm \sigma)} \qquad \text{R-boundary condition}$$
$$\psi_{\pm}^{\mu}(\tau \pm \sigma) = \frac{1}{\sqrt{2}} \sum_{r \in \mathbb{Z} + 1/2} \psi_{r}^{\mu} e^{-ir(\tau \pm \sigma)} \qquad \text{NS-boundary condition}$$

The Majorana condition requires

$$\psi^{\mu}_{-r} = (\psi^{\mu}_{r})^{\star}$$

Note: Only the R sector gives a zero mode ψ_0^{μ} (that should lead to fermions).

Boundary conditions and mode expansion: closed strings

For closed strings the surface terms vanish when the boundary conditions are periodic or anti-periodic, separately, for each component of ψ^{μ}_{\pm} ,

$$\psi^{\mu}_{-}(\tau-\sigma) = \sum_{r} \psi^{\mu}_{r} \epsilon^{-2ir(\tau-\sigma)} \qquad \psi^{\mu}_{+}(\tau+\sigma) = \sum_{r} \bar{\psi}^{\mu}_{r} \epsilon^{-2ir(\tau+\sigma)}$$

with the mode index *r* constrained according to either R or NS conditions. The bar on $\bar{\psi}_r^{\mu}$ shouldn't be confused with the one in Dirac's equation.

Corresponding to the different pairings between left- and right-moving modes, there are four distinct closed string sectors that can be grouped according to the space-time nature of their quantum states:

R - R and NS - NS (bosons) R - NS and NS - R (fermions)

The superstring has, as we will shortly prove, space-time fermions!

Not surprisingly, they come from the *odd* combination of boundary conditions.

Super Virasoro constraints: closed strings

The closed string mode expansion of the physical constraints are given by

$$T_{--}(\xi^{-}) = \alpha' \sum_{n=-\infty}^{\infty} L_n \, \epsilon^{-2i\,n\,\xi^{-}} \qquad T_{++}(\xi^{+}) = \alpha' \sum_{n=-\infty}^{\infty} \bar{L}_n \, \epsilon^{-2i\,n\,\xi^{+}}$$

as well as for the current constraints:

$$J_{-}(\xi^{-}) = \alpha' \sum_{r} G_{r} \epsilon^{-2ir\xi^{-}} \qquad J_{+}(\xi^{+}) = \alpha' \sum_{r} \overline{G}_{r} \epsilon^{-2ir\xi^{+}}$$

where, inverting the above expressions, the super-Virasoro generators read

$$L_{n} = \frac{1}{2} \sum_{m=-\infty}^{\infty} \alpha_{n-m} \cdot \alpha_{m} + \frac{1}{4} \sum_{r} (2r - n) \psi_{n-r} \cdot \psi_{r}$$
$$G_{r} = \sum_{m=-\infty}^{\infty} \alpha_{m} \cdot \psi_{r-m} \qquad (\text{analogous for } \bar{L}_{n} \text{ and } \bar{G}_{r})$$

These expressions result from mode expansion of the world-sheet fields.

Normal ordering and zero-point energy

Going quantum, we have to be careful about normal ordering of operators:

$$L_n = \frac{1}{2} \sum_{m=-\infty}^{\infty} : \alpha_{n-m} \cdot \alpha_m : + \frac{1}{4} \sum_{r} (2r-n) : \psi_{n-r} \cdot \psi_r :$$

an ambiguity arising for L_0 (and \overline{L}_0).

As before, the constraints must be imposed through a physical state condition

$$L_{m>0} |\mathrm{phys}
angle = 0$$
 $G_{r>0} |\mathrm{phys}
angle = 0$ $(L_0 - a) |\mathrm{phys}
angle = 0$

(same for \overline{L}_m and \overline{G}_r) for a given *c*-number *a* that we shall determine.

This computation, as for the bosonic string, is better performed in space-time light-cone coordinates

$$x^{\pm} := \frac{1}{\sqrt{2}} (x^0 \pm x^{D-1}) \qquad \psi^{\pm} := \frac{1}{\sqrt{2}} (\psi^0 \pm \psi^{D-1})$$

while x^i and ψ^i are the remaining space-like coordinates of the superstring.

Light-cone gauge

Recall that in the bosonic string, as we will do now, we gauged away α_n^+ by choosing the light-cone gauge $\tau \sim x^+$. We shall take at the same time $\psi^+ = 0$ if we want to be consistent with world-sheet supersymmetry.

Given that $\partial_+ x^+ = \frac{1}{2}p^+$ and $\psi^+ = 0$, we can solve for x^- and ψ^-

$$\partial_{+}x^{-} = \frac{1}{\alpha' p^{+}} \left(\partial_{+}x^{i} \partial_{+}x^{i} + \frac{i}{2}\psi^{i} \partial_{+}\psi^{i} \right) \qquad \psi^{-} = \frac{2}{\alpha' p^{+}}\psi^{i} \partial_{+}x^{i}$$

in terms of the space-like components. This implies for the oscillators

$$\alpha_n^- = \frac{1}{\sqrt{2\alpha'}\rho^+} \left[\sum_{m=-\infty}^{\infty} : \alpha_{n-m}^i \alpha_m^i : + \frac{1}{2} \sum_r (2r-n) : \psi_{n-r}^i \psi_r^i : -2a \,\delta_n \right]$$

$$\psi_r^- = \sum_{m=-\infty}^{\infty} \alpha_m^i \psi_{r-m}^i \qquad (\text{analogous for } \bar{\alpha}_n^- \text{ and } \bar{\psi}_r^-)$$

The same computations can be performed in the case of open strings.

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Super Virasoro algebra

By relying in the commutation relation of the x^{μ} and the anti-commutators

 $\{\psi_{r}^{\mu},\,\psi_{s}^{\nu}\}=\delta_{r+s,0}\,\eta^{\mu\nu}$

it can be shown that L_n and G_r generate the $\mathcal{N} = 1$ supersymmetric extension of the Virasoro algebra:

$$[L_n, L_m] = (n - m) L_{n+m} + \frac{c}{12} A(n) \delta_{n+m,0}$$

$$[L_n, G_r] = \frac{1}{2} (n - 2r) G_{n+r} \Rightarrow Osp(1|2) \text{ subalgebra for NS}$$

$$\{G_r, G_s\} = 2 L_{r+s} + \frac{c}{3} B(r) \delta_{r+s,0}$$

$$A(n) = n^3 - n \text{ and } B(r) = r^2 - 1/4 \text{ for the NS sector,}$$

$$A(n) = n^3 \text{ and } B(r) = r^2 \text{ for the R sector.}$$

 $c = D + \frac{D}{2}$ is the total contribution to the conformal anomaly.

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Super Virasoro algebra: the conserved angular momentum

The densities of momentum and angular momentum along the string are Noether currents

$$P^{\mu}_{\alpha} = \partial_{\alpha} x^{\mu} \qquad J^{\mu\nu}_{\alpha} = x^{\mu} \partial_{\alpha} x^{\nu} - x^{\nu} \partial_{\alpha} x^{\mu} + i \bar{\psi}^{\mu} \rho_{\alpha} \psi^{\nu}$$

Inserting the mode expansions for \mathbf{x}^{μ} and ψ^{μ}

$$J^{\mu\nu} = \int_0^{\pi} J^{\mu\nu}_{\tau} d\sigma = J^{\mu\nu}_B + J^{\mu\nu}_F = \ell^{\mu\nu} + E^{\mu\nu} + J^{\mu\nu}_F$$

where

$$\ell^{\mu\nu} = x^{\mu} p^{\nu} - x^{\nu} p^{\mu} \qquad E^{\mu\nu} = -i \sum_{n=1}^{\infty} \frac{1}{n} \left(\alpha^{\mu}_{-n} \alpha^{\nu}_{n} - \alpha^{\nu}_{-n} \alpha^{\mu}_{n} \right)$$

$$J_{F}^{\mu\nu} = -i \sum_{r \ge 0} \left(\psi_{-r}^{\mu} \, \psi_{r}^{\nu} - \psi_{-r}^{\nu} \, \psi_{r}^{\mu} \right)$$
 NS

$$J_{F}^{\mu\nu} = -\frac{i}{2} [\psi_{0}^{\mu}, \psi_{0}^{\nu}] - i \sum_{k=1}^{\infty} \left(\psi_{-k}^{\mu} \psi_{k}^{\nu} - \psi_{-k}^{\nu} \psi_{k}^{\mu} \right) \qquad \mathbf{R}$$

Mass spectrum of the open string

As in the bosonic string case, we can define a number operator as

$$N := \sum_{m=1}^{\infty} \alpha_{-m}^{i} \alpha_{m}^{i} + \sum_{r>0} r \psi_{-r}^{i} \psi_{r}^{i}$$
$$L_{0} = -\alpha' M^{2} + N \qquad \Rightarrow \qquad M^{2} = \frac{1}{\alpha'} (n - a)$$

since $(L_0 - a) | \text{phys} \rangle = 0$ and $N | \text{phys} \rangle = n | \text{phys} \rangle$.

The open string spectrum of states has two independent NS and R sectors. Let us start by analyzing the NS sector.

The NS Fock-space unique nondegenerate ground state $|k; 0\rangle_{NS}$ satisfies

$$lpha_{m>0}^{\mu}|k;0
angle_{
m NS}=\psi_{r>0}^{\mu}|k;0
angle_{
m NS}=0$$
 $lpha_{0}^{\mu}|k;0
angle_{
m NS}=\sqrt{2lpha'}\,k^{\mu}|k;0
angle_{
m NS}$

Since $N |k; 0\rangle_{NS} = 0$, this vacuum state would be tachyonic for a > 0:

$$M^2 = -\frac{a}{\alpha'}$$
 as for the bosonic string!

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Determination of the normal ordering constant

In order to evaluate \underline{a} , we must reorder \underline{L}_0 :

$$L_{0} = \frac{1}{2} \sum_{m=-\infty}^{\infty} \alpha_{-m}^{i} \alpha_{m}^{i} + \frac{1}{2} \sum_{r=-\infty}^{\infty} \left(r + \frac{1}{2}\right) \psi_{-r-1/2}^{i} \psi_{r+1/2}^{i}$$

Taking into account that $[\alpha_m^i, \alpha_n^j] = m \delta^{ij} \delta_{mn}$ and $\{\psi_{-s}^i, \psi_t^j\} = \delta^{ij} \delta_{st}$,

$$L_0 = -\alpha' M^2 + N + \frac{D-2}{2} \sum_{m=1}^{\infty} m - \frac{D-2}{2} \sum_{r=0}^{\infty} \left(r + \frac{1}{2}\right)$$

Thus, we conclude that, since $(L_0 - a) |phys\rangle = 0$,

$$a_{\rm NS} = -\frac{D-2}{2} \left[\sum_{m=1}^{\infty} m - \sum_{r=0}^{\infty} \left(r + \frac{1}{2} \right) \right]$$

in the NS sector. This expression demands regularization.

However, we can realize at this point that $a_{R} = 0$ (the magic of SUSY!).

Regularized infinite sum

It is useful to compute in general (for $\alpha > 0$; for the bosonic string, $\alpha = 0$),

$$Z_{\alpha} = \frac{1}{2} \sum_{n=0}^{\infty} (n+\alpha) = -\frac{1}{24} + \frac{1}{4}\alpha(1-\alpha)$$

as seen from the $\epsilon \rightarrow 0$ limit of the finite part of

$$\begin{aligned} Z_{\alpha}(\epsilon) &= \frac{1}{2} \sum_{n=0}^{\infty} \left(n+\alpha \right) e^{-(n+\alpha)\epsilon} = -\frac{1}{2} \frac{d}{d\epsilon} \sum_{n=0}^{\infty} e^{-(n+\alpha)\epsilon} = -\frac{1}{2} \frac{d}{d\epsilon} \left(\frac{e^{-\alpha\epsilon}}{1-e^{-\epsilon}} \right) \\ &= -\frac{1}{2} \frac{d}{d\epsilon} \left[\left(1-\alpha\epsilon + \frac{1}{2}\alpha^{2}\epsilon^{2} + \mathcal{O}(\epsilon^{3}) \right) \left(\frac{1}{\epsilon} + \frac{1}{2} + \frac{1}{12}\epsilon + \mathcal{O}(\epsilon^{2}) \right) \right] \\ &= -\frac{1}{2} \frac{d}{d\epsilon} \left[\frac{1}{\epsilon} + \frac{1}{2} (1-2\alpha) + \frac{1}{12} (1-6\alpha + 6\alpha^{2})\epsilon + \mathcal{O}(\epsilon^{2}) \right] \\ &= \frac{1}{2\epsilon^{2}} - \frac{1}{24} + \frac{1}{4}\alpha(1-\alpha) + \mathcal{O}(\epsilon) \end{aligned}$$

The critical dimension

With this regularization we can confirm that $a_{\rm R} = 0$, and

$$a_{\rm NS} = (D-2)(Z_{1/2}-Z_0) = \frac{D-2}{16} \Rightarrow a_{\rm NS} = \frac{1}{2}$$

The tachyonic vacuum of the NS sector see also GSW, end of § 4.2.2 and § 4.3.1

$$M^2 = -\frac{D-2}{16\,\alpha'}$$

The first excited level can be constructed by acting with $\psi_{-\frac{1}{4}}^{i}$ on $|\mathbf{k}; \mathbf{0}\rangle_{\rm NS}$,

$$N\left(\psi_{-\frac{1}{2}}^{i}|\boldsymbol{k};\boldsymbol{0}\rangle_{\mathrm{NS}}\right) = \frac{1}{2}\left(\psi_{-\frac{1}{2}}^{i}|\boldsymbol{k};\boldsymbol{0}\rangle_{\mathrm{NS}}\right)$$

Thus, its mass can be readily determined to be

$$M^2 = \frac{1}{2 \alpha'} \left(1 - \frac{D-2}{8} \right) \qquad \Rightarrow \qquad D = 10$$

This vector has D - 2 components; in a relativistic theory it must be massless (describe the physical polarizations of an open string photon field $A_{\mu}(x)$).

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The open string spectrum

The NS sector of the RNS superstring is quite reminiscent of the bosonic string spectrum.

Indeed, all NS states are space-time bosons: they transform in appropriate irreducible representations of SO(8), which is the Little group of SO(1,9).

In the R sector there are zero modes ψ_0^{μ} which satisfy the 10d Dirac algebra!

 $\left\{\psi_0^\mu\,,\,\psi_0^\nu\right\}=\eta^{\mu\nu}$

As announced, the zero modes can be regarded as Dirac matrices,

 $\psi_0^{\mu} = \frac{1}{\sqrt{2}} \Gamma^{\mu} \qquad \Rightarrow \qquad \text{degenerate ground state}$

This is the origin of space-time fermions in the superstring.

In particular they are finite dimensional operators: all states in the R sector should be space-time fermions in order to furnish representation spaces on which these operators can act.